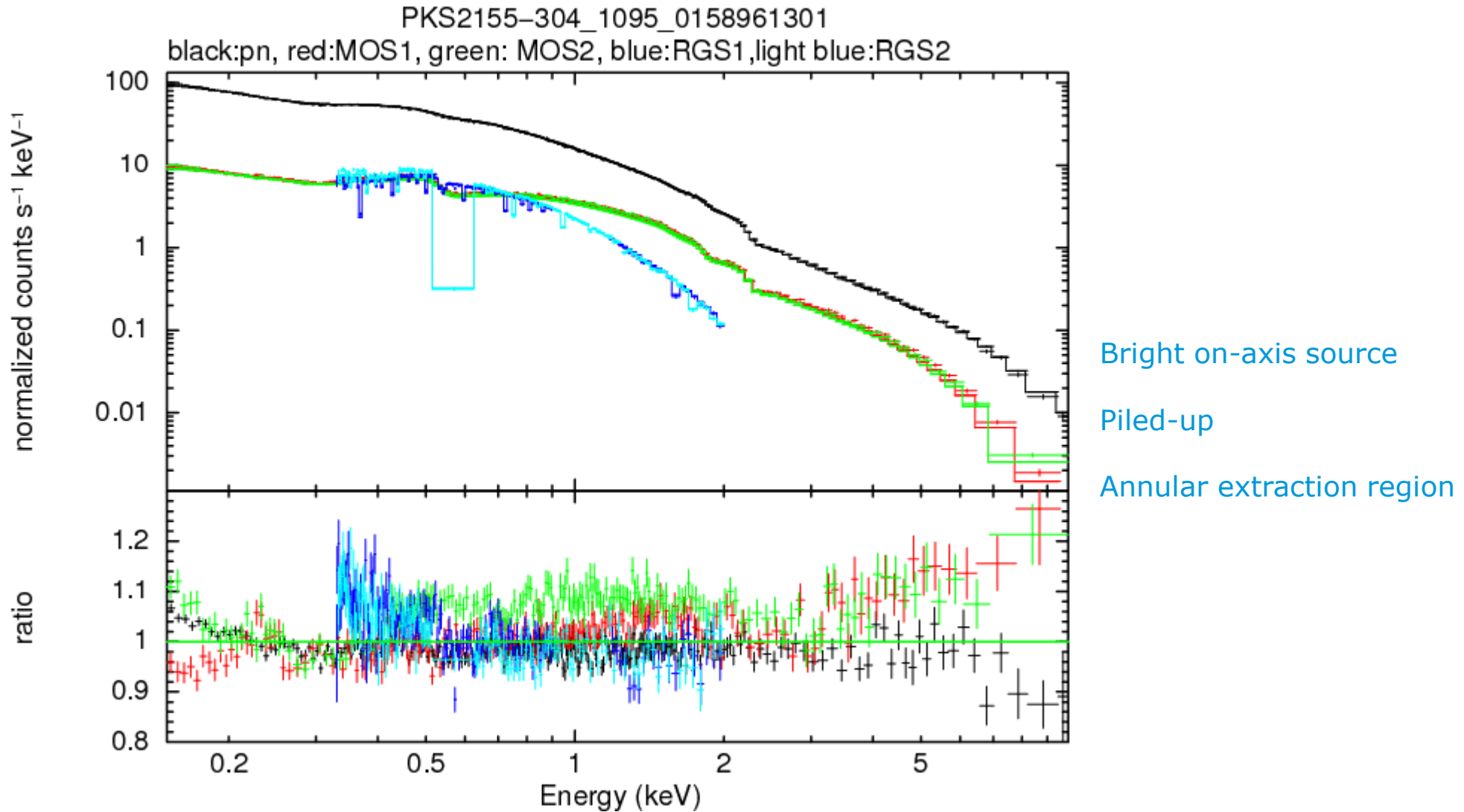
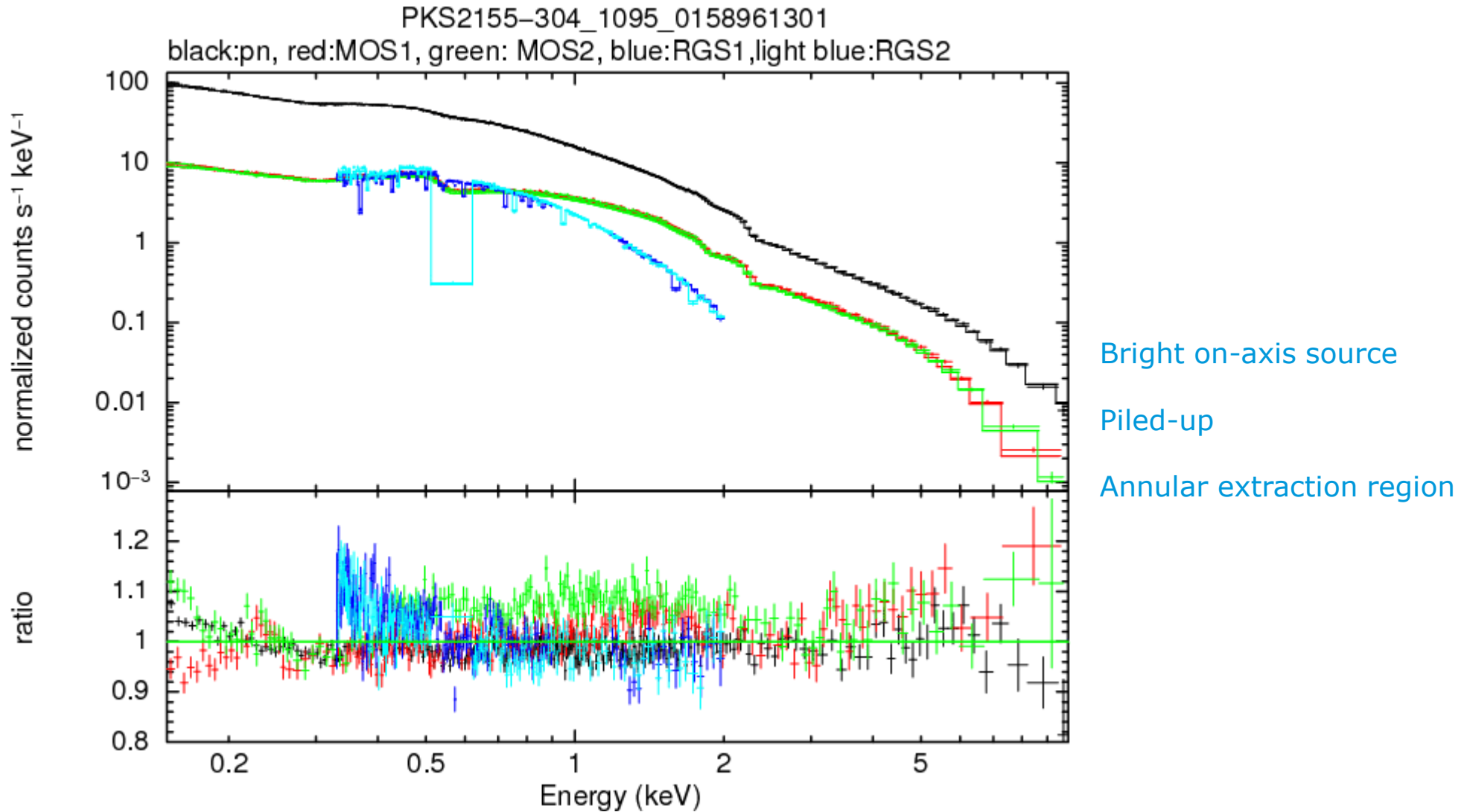


Investigation of the XRT XPSF Wings

Michael Smith, ESAC

EPIC Cal Ops Meeting, ESAC, 9-10 April 2015





XRT3_XPSF_0016.CCF:

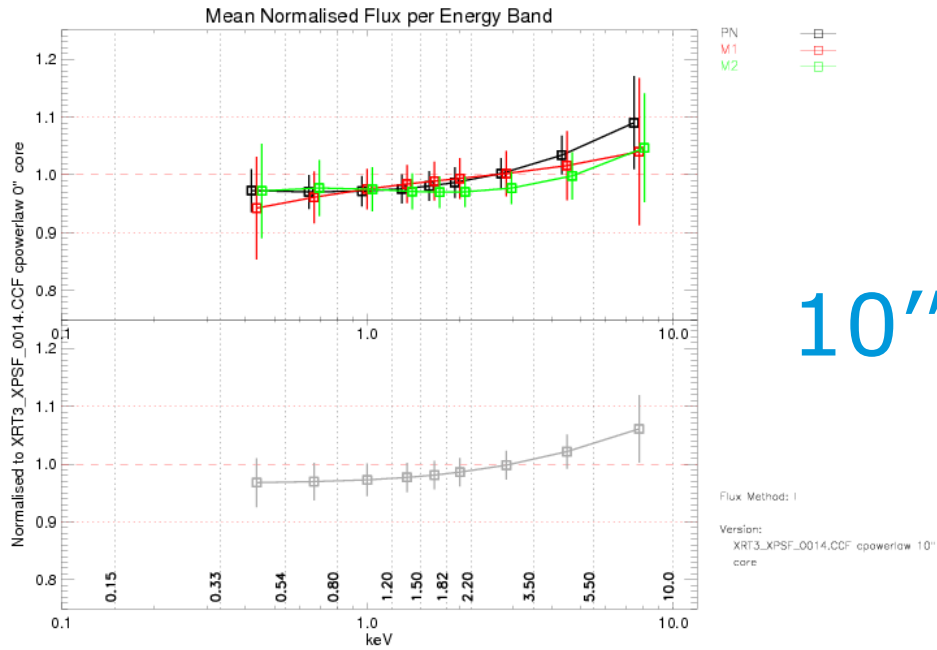
- Refinement of the high-energy XRT3 (=PN) PSF
- Based on TI mode data (Mkn 421)
- Optimise parameters describing PSF wings in order to obtain constant measured flux independent of number of excised columns

0014 versus 0016 on Imaging Mode Data

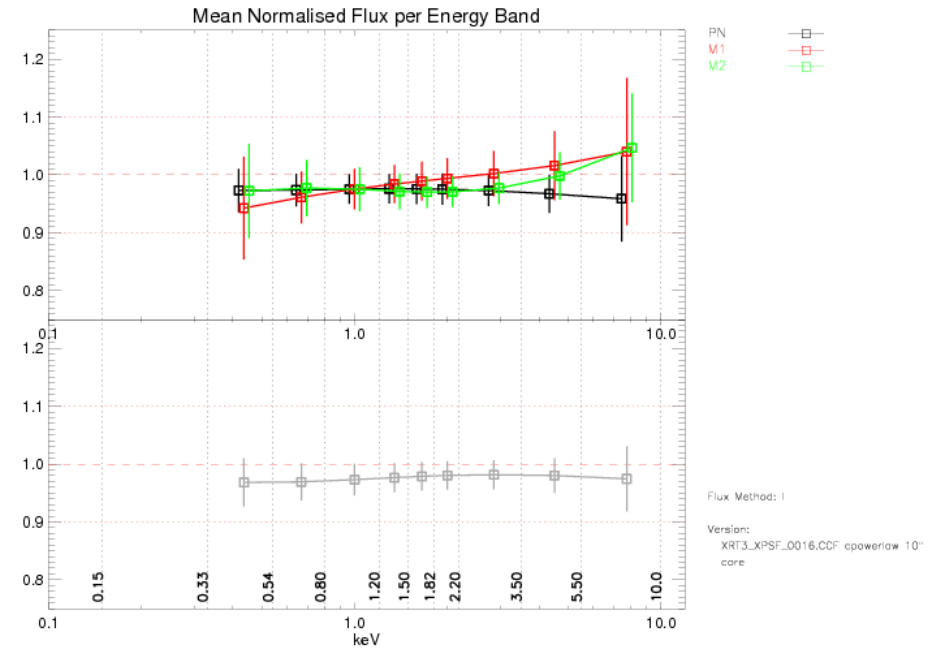


Sample of ~ 30 on-axis non-piled-up sources
Determine band fluxes of various annular regions (varying inner radius)
normalised to those of a circular region.

XRT_XPSF_0014.CCF



XRT_XPSF_0016.CCF

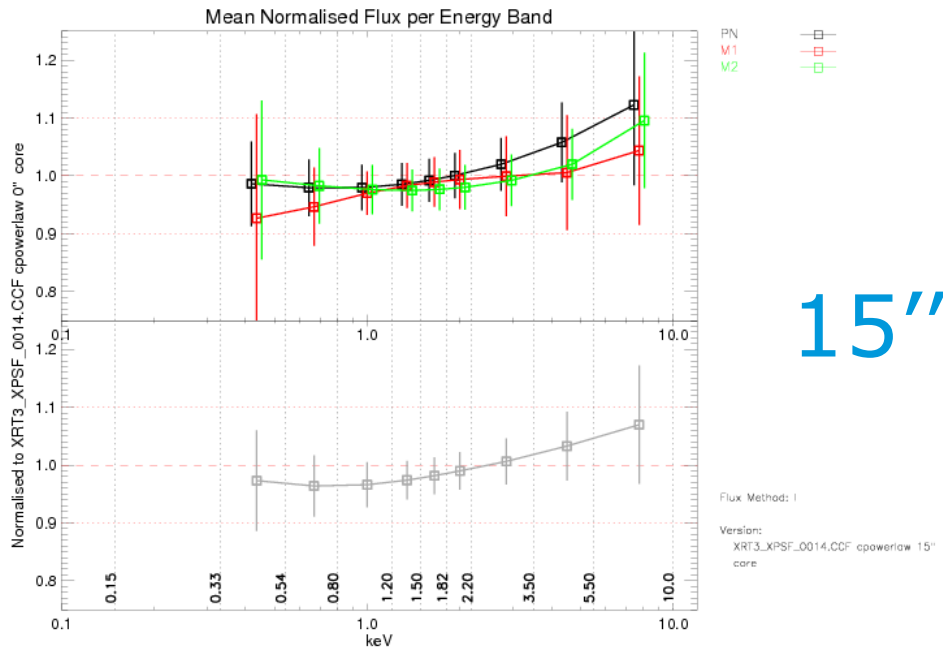


0014 versus 0016 on Imaging Mode Data

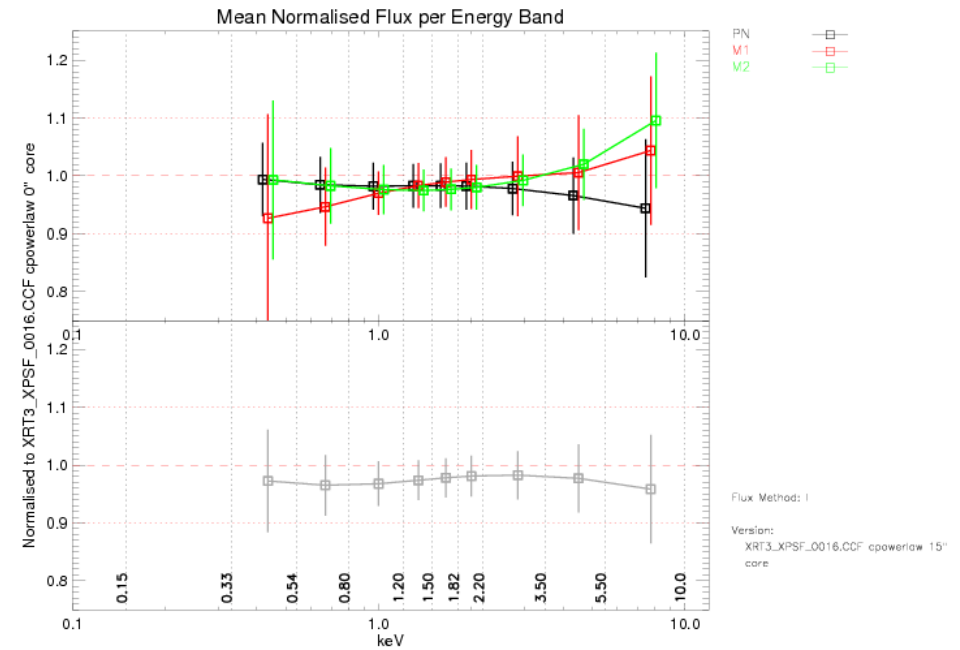


Sample of ~ 30 on-axis non-piled-up sources
Determine band fluxes of various annular regions (varying inner radius)
normalised to those of a circular region.

XRT_XPSF_0014.CCF



XRT_XPSF_0016.CCF

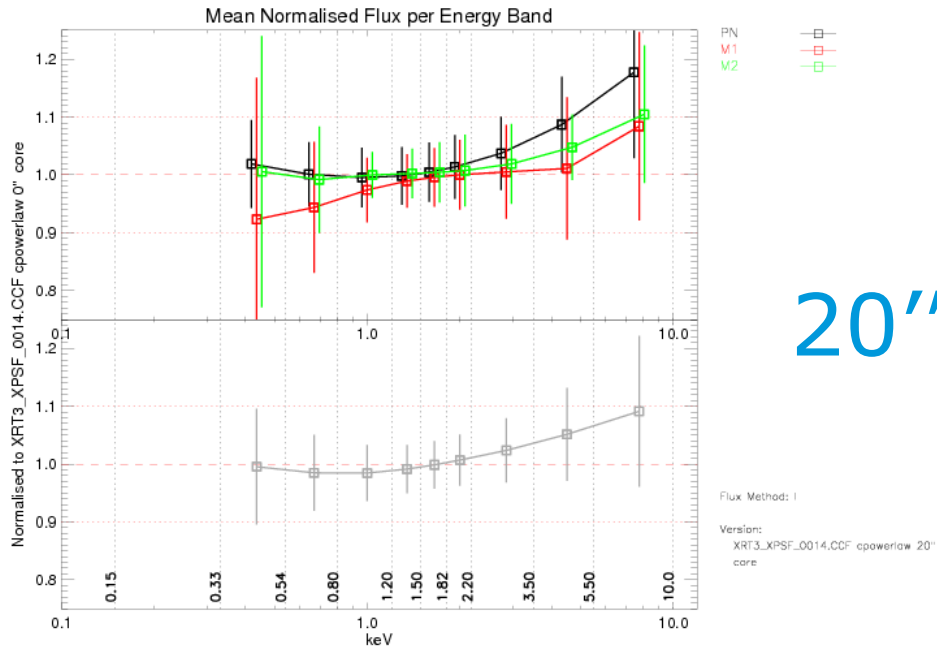


0014 versus 0016 on Imaging Mode Data

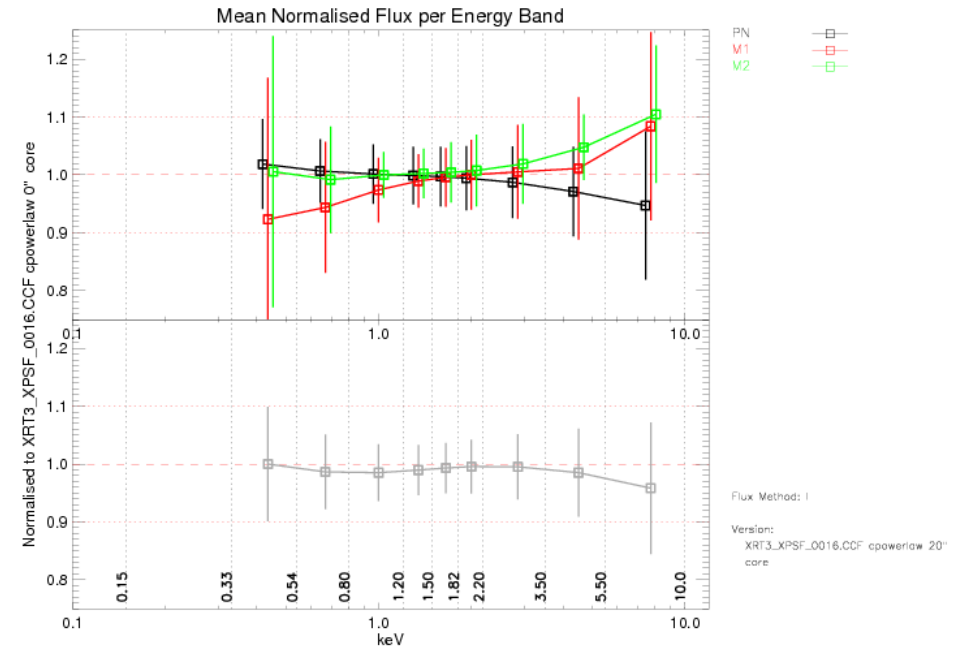


Sample of ~ 30 on-axis non-piled-up sources
Determine band fluxes of various annular regions (varying inner radius)
normalised to those of a circular region.

XRT_XPSF_0014.CCF



XRT_XPSF_0016.CCF



XRT_XPSF_0016.CCF does what is advertised for PN

Do the MOS XRT PSFs need a similar modification?

Test this, using the same method as used for deriving 0016

But on MOS (and PN) imaging mode data

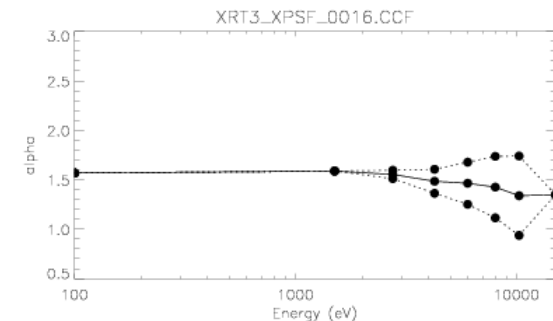
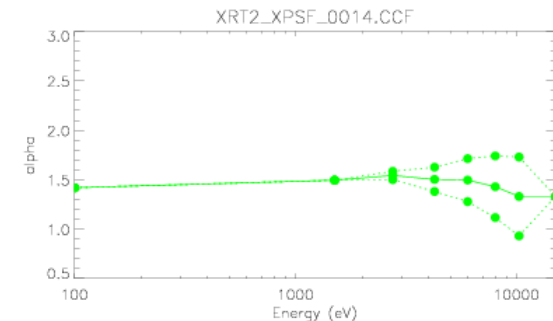
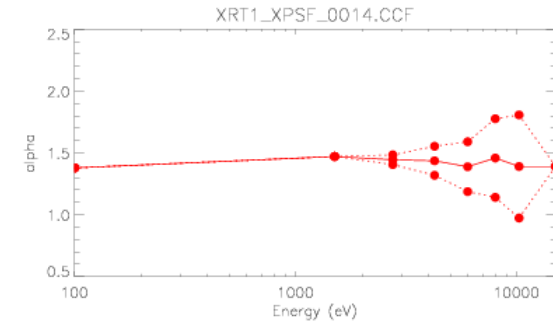
Using a sample of ~ 30 on-axis non-piled-up sources

Sample PSF parameter space within reasonable limits.

5 PSF parameters(E):

- beta2d core radius, r0
- beta2d power-law slope, alpha
- ellipticity, epsilon
- gauss2d FWHM
- gauss2d-to-beta2d normalisation

XRT3_XPSF_0016 essentially differed in alpha:
0% at 2keV up to -16% change at 10.5 keV.



For each:

- alpha parameter value, and
 - annular inner radius
- determine the mean band fluxes.

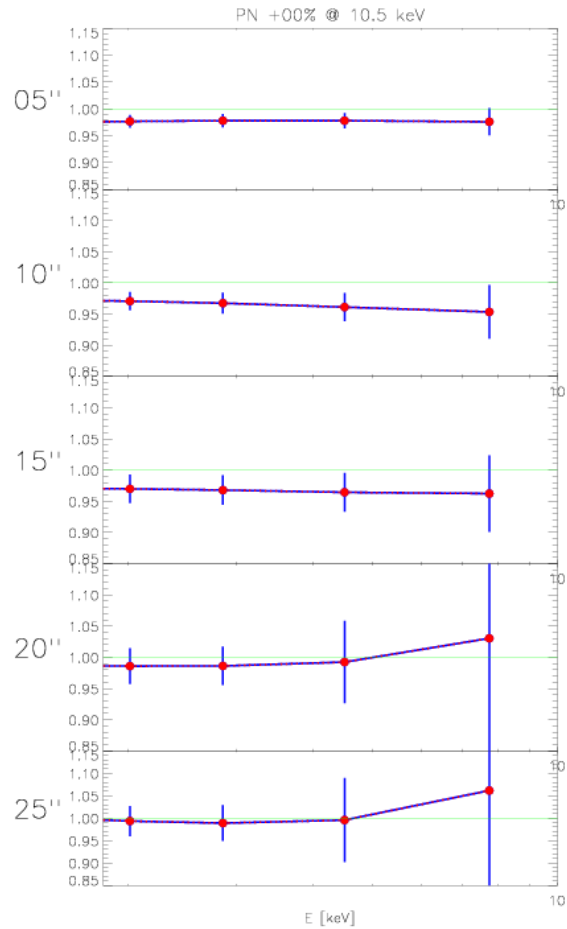
Normalise to those of a circular extraction.

Determine the parameter value for which the energy dependence of the flux ratios is minimal, considering the set of annular inner radii.

Results for alpha

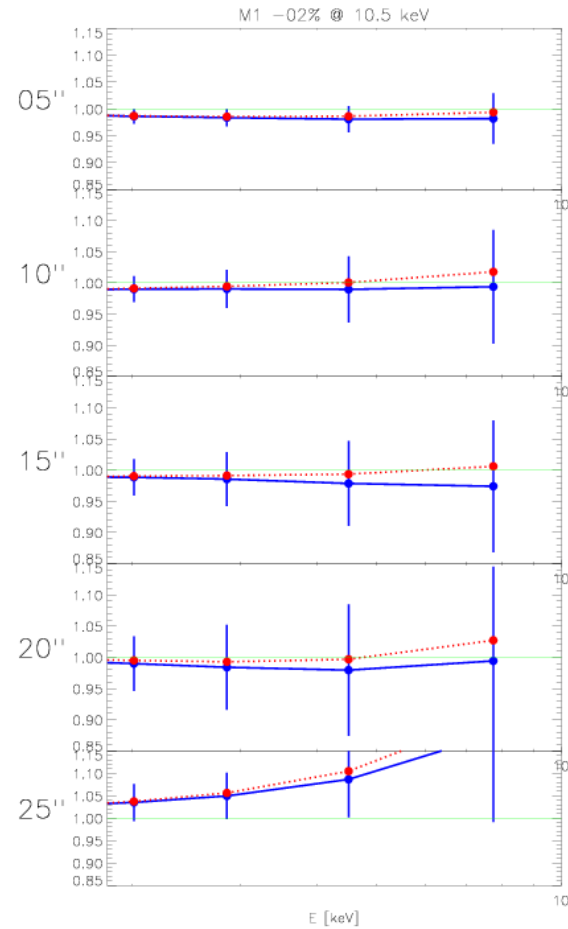
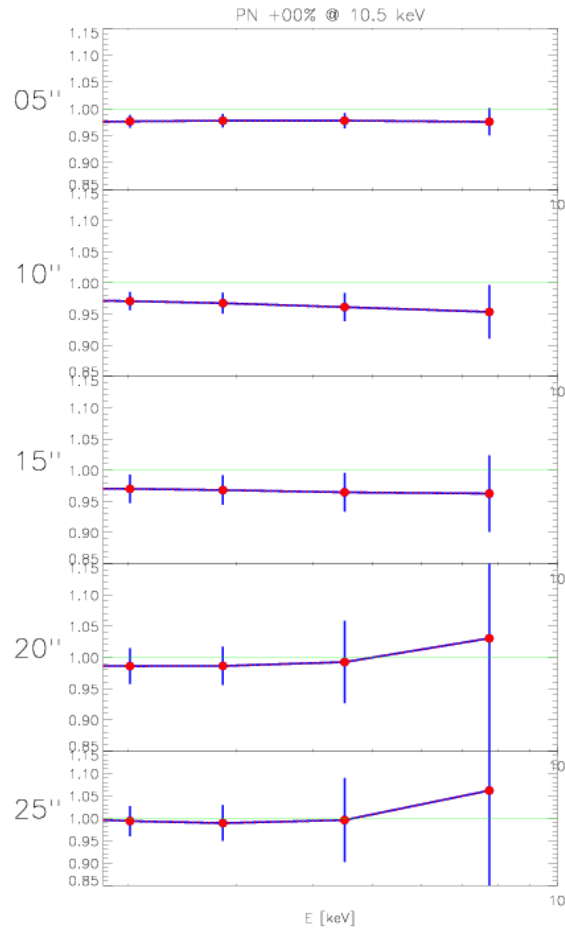


PN: ~0%



PN: ~0%

M1: ~0%



Results for alpha



PN: $\sim 0\%$

M1: $\sim 0\%$

M2: -10%

