

# RGS – EPIC-pn Cross Calibration

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## RGS Calibration Status

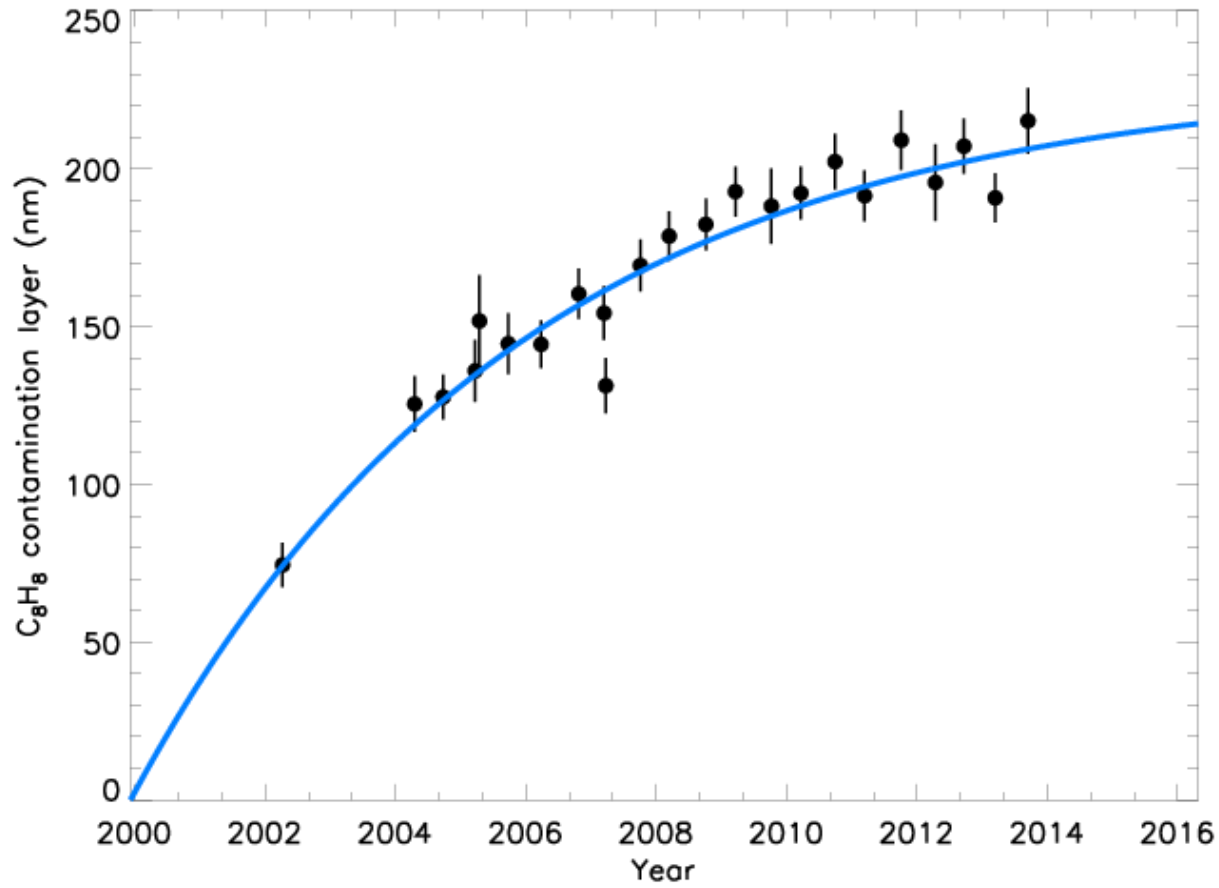
Wavelength scale accuracy	1 <sup>st</sup> order: 6 mÅ 2 <sup>nd</sup> order: 5 mÅ no systematic offset
Line Spread Function	$\Delta \text{FWHM}/\text{FWHM} \leq 10\%$
Effective area	Absolute effective area: 10 %  Bad (hot/cool) pixels: less than 3% of detectors surface. No changes in the last years  Contamination: 47% decrease in effective area at 35 Å, taken into account in the calibration. Same for RGS1 and RGS2, within errors.

For details, refer to

“Calibration and in-orbit performance of the RGS on board XMM”, de Vries et al., 2015, A&A 573, p128

“Status of the RGS Calibration”, XMM-SOC-CAL-TN-0030, July 2014

## RGS Contamination model



From RXJ1856 countrate at 35 Å

Assuming that the contaminant is a type of Hydrocarbon (C<sub>8</sub>H<sub>8</sub>) coming from the carbon-fiber structures of the telescope tube.

## RGS/EPIC-pn rectification factors (I)

### Motivation:

Indications of systematic differences in flux from EPIC-pn and RGS longward 24 Å  
( $E < 0.5$  keV)

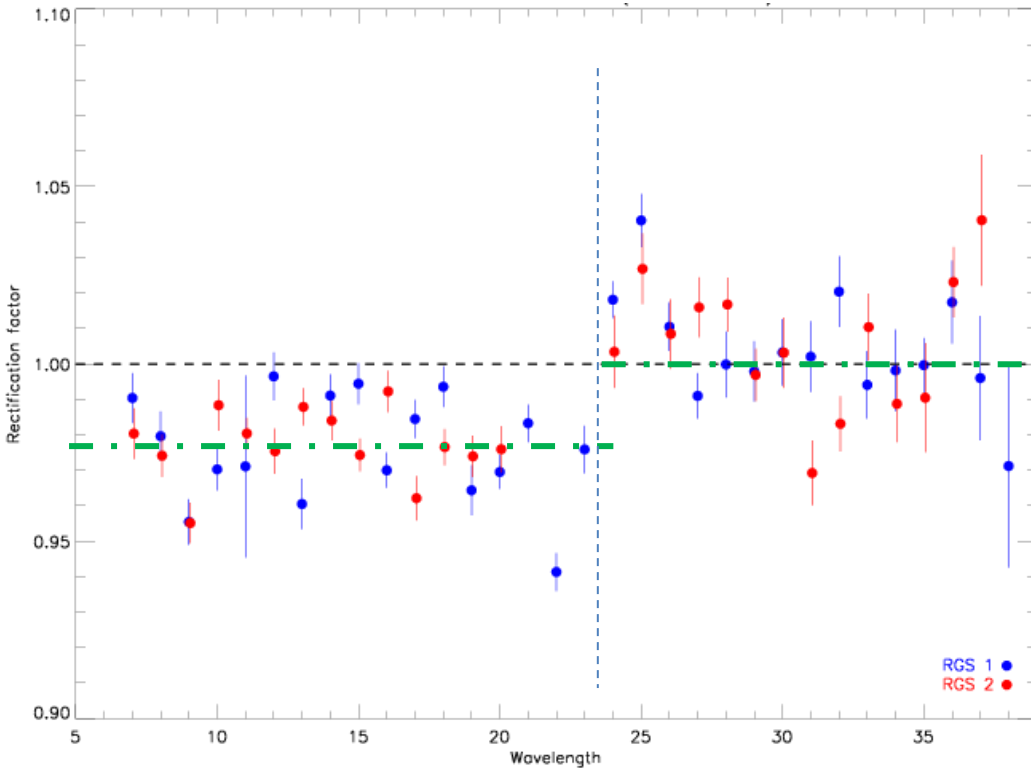
Derived from 43 observations of 3C 273, PKS 2155-304 and H 1426+428 processed with SASv10.  
All EPIC-pn Small Window mode

Flux ratio (RGS[12]/EPIC-pn) computed in 1 Å step through XSPEC fits:

- Spectra binned to 25 cts/bin
- Chisq statistics
- Fixed  $N_H$
- Model:
  - PKS 2155 and H 1426: Broken power law
  - 3C 273: Two power laws

see XMM-CCF-REL-269, XMM-SOC-CAL-TN-0089

## RGS/EPIC-pn rectification factors (II)



- Implemented as an step function:

	7 - 23.5 Å	23.5 - 38.5 Å
<b>RGS 1</b>	<b><math>0.97 \pm 0.02</math></b>	<b><math>1.00 \pm 0.01</math></b>
<b>RGS 2</b>	<b><math>0.98 \pm 0.01</math></b>	<b><math>1.00 \pm 0.02</math></b>

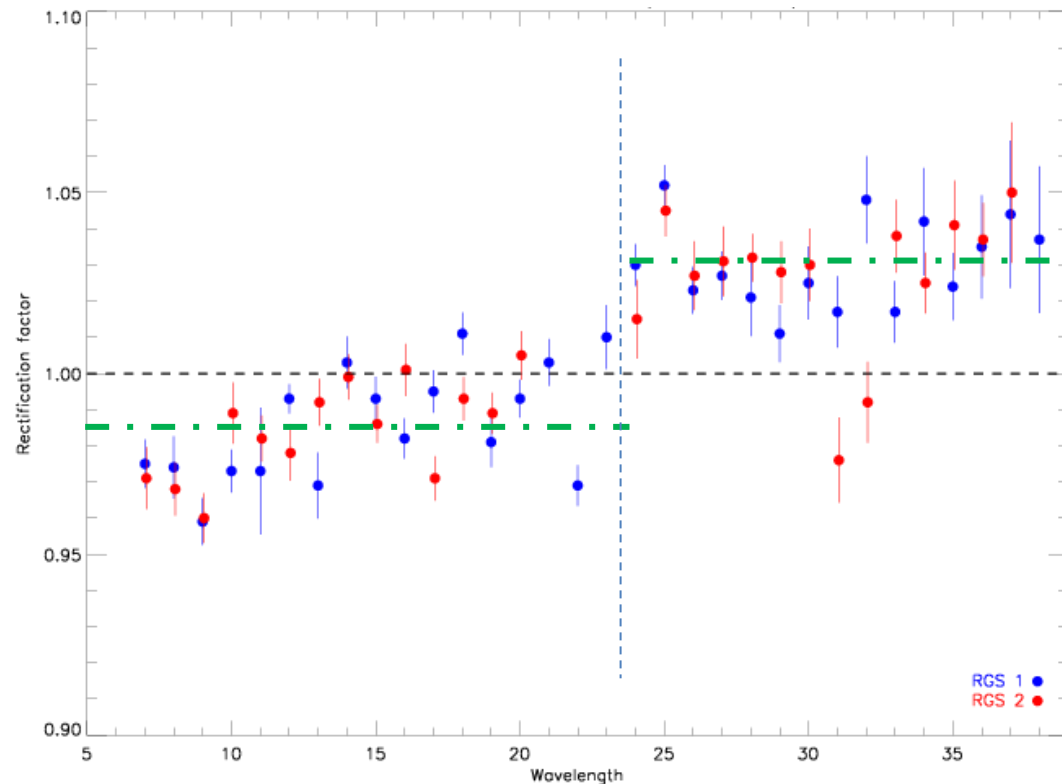
- Factors stored in extension `RECTIFICATION` in `RGS[12]_EFFAREACORR.CCF`
- Correction applied to the RGS effective area via the `rgsrmfgen` parameter  
`withrectification=yes`

## Need for new rectification factors ?

- More data (51 observations)
- EPIC-pn and RGS processed with SASv14 and calibrations available in November 2014

Differences from previous may come from:

- Different sample
- New RGS contamination model
- New EPIC-pn calibrations (PSF, CTI, redistribution...)
- Different extraction regions for EPIC-pn spectra (larger inner annuli radius, more accurate pile-up correction)



	7 - 23.5 Å	23.5 - 38.5 Å
<b>RGS 1</b>	<b>0.98 ± 0.02</b>	<b>1.03 ± 0.01</b>
<b>RGS 2</b>	<b>0.99 ± 0.01</b>	<b>1.03 ± 0.02</b>

Need to be re-evaluated every time calibrations and/or SW change