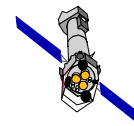


MOS CTI correction

Bruno Altieri

EPIC calibration meeting

Milano, 6-8 November 2001

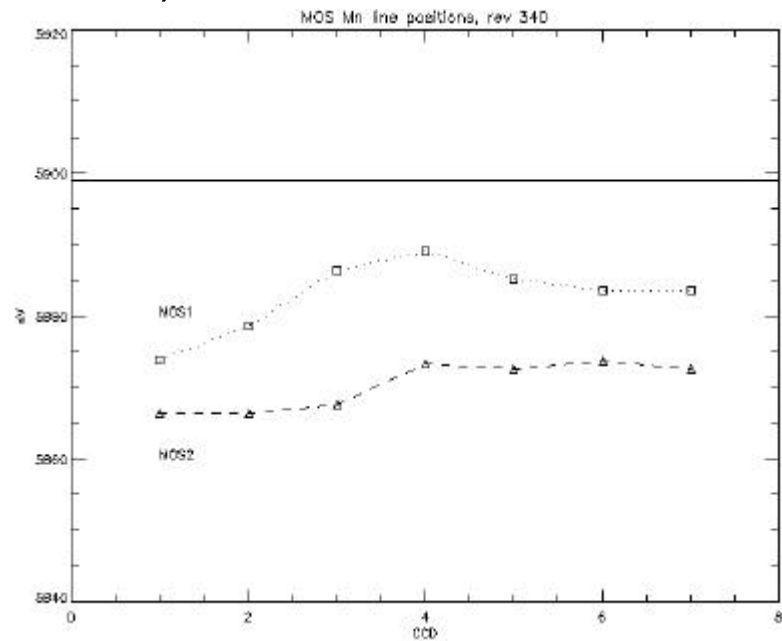


XMM

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Current MOS CTI

- In current SAS (v5.2, MOS CTI CCF v5) the CTI is under-corrected ==> line centroid is shifted to lower energies
- Discrepancy increasing with time (rev #)
 - By now (rev 340) it is about **25-30 eV** at Mn energies

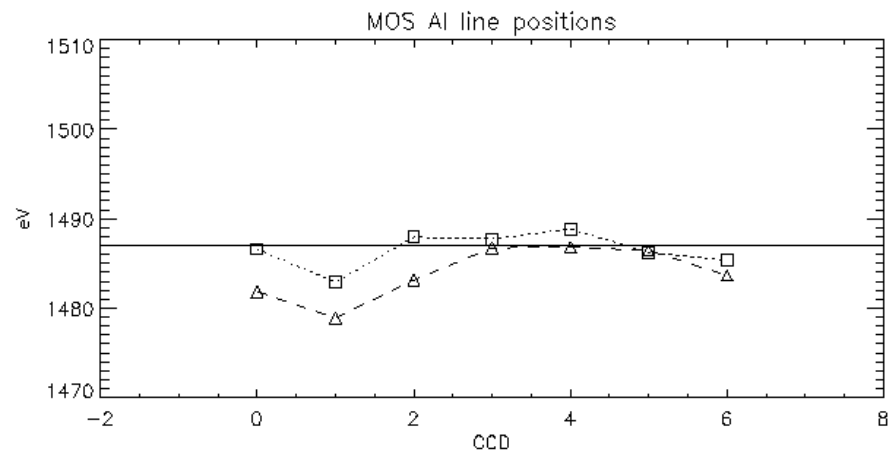
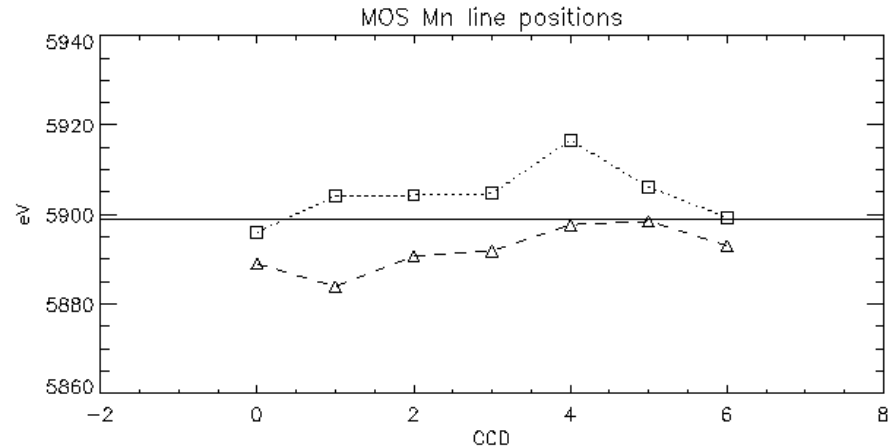


MOS CTI algorithm

- X = serial loss, Y: parallel loss
- $CTI = CTIX \cdot RAWX + CTIY \cdot RAWY$
- **Current CTI correction :**
 - $CTIX = \delta T \cdot rate_x + (a_{0xCCDi} + b_{0xCCDi} \cdot PHA)$
 - $CTIY = \delta T \cdot rate_y \cdot (PHA)^{1/2} + (a_{0yCCDi} + b_{0yCCDi} \cdot PHA)$
 - with $(a,b)_{0X,Y}$ from ground-based test and $\delta T = T - T_0$
- **New CTI correction:**
 - $CTIX = a_{1xCCDi} + b_{1xCCDi} \cdot PHA$
 - $CTIY = \delta T \cdot rate_{yCCDi} \cdot (PHA)^{\alpha_{CCDi}} + (a_{1yCCDi} + b_{1yCCDi} \cdot PHA)$
 - with $(a,b)_{1X,Y}$ from ground-based test and $\delta T = T - T_{launch}$
 - 6 parameters per CCD (4 for parallel loss and 2 for serial)
 - degradation rate and power index (0.55 to 0.7) CCD dependant !
 - All parameters derived from a linear fit of MOS CTI plots.

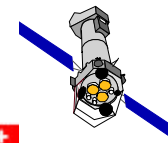
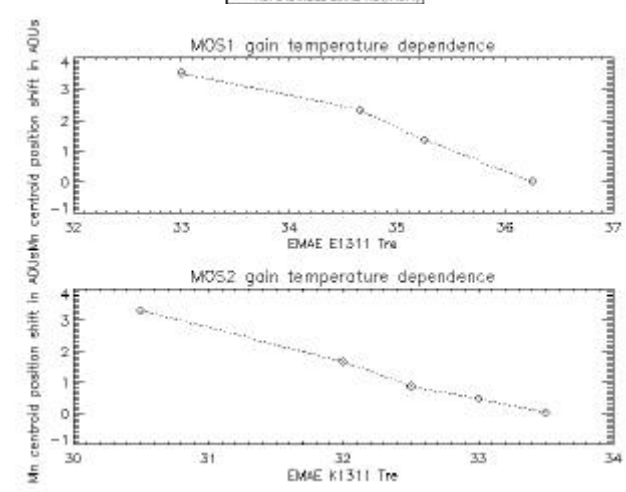
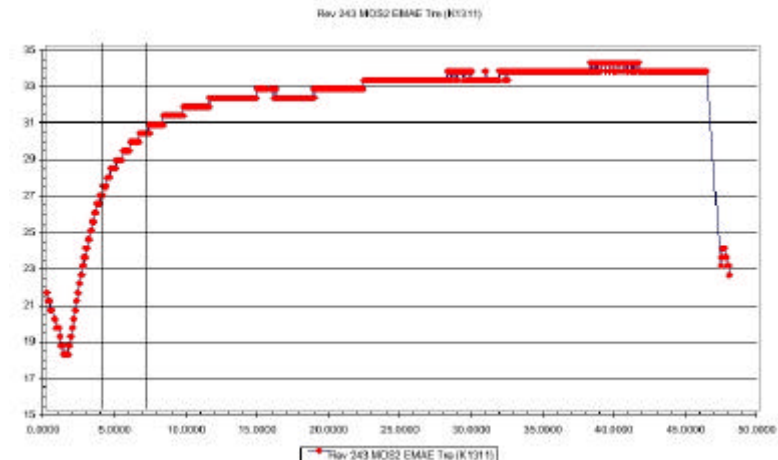
MOS CTI new algo tests

- **New algorithm tested in:**
 - CCF MOS CTI **version 6**
 - combined with **cal-3.120**
- **Some slight over-correction for MOS1**
- **Still some under-correction for MOS2**
 - CTI underestimated because of high MOS2 calibration source ?
- **Proposed for SAS v5.3**



Temperature-dependence of the gain ?

- CTI correction does not work for cal-closed during the eclipse seasons, because of the EMAE T^{re} variations.
- EMAE T^{re} colder \implies CTI is less \implies over-correction in SAS.
- Effect worse for MOS2 than MOS1 and energy-dependent. $dPHA = a \cdot E/K1311 + b + c \cdot PHA$
- From rev 242 long calclosed : $dPHA / dT^{re} \sim 1$
- Effect worse in autumn eclipses, as T^{re} excursions are larger.
- Science observation less affected, but T^{re} stabilized at ~ 2 degrees lower in last eclipse season, $\implies \sim 7$ eV shift.

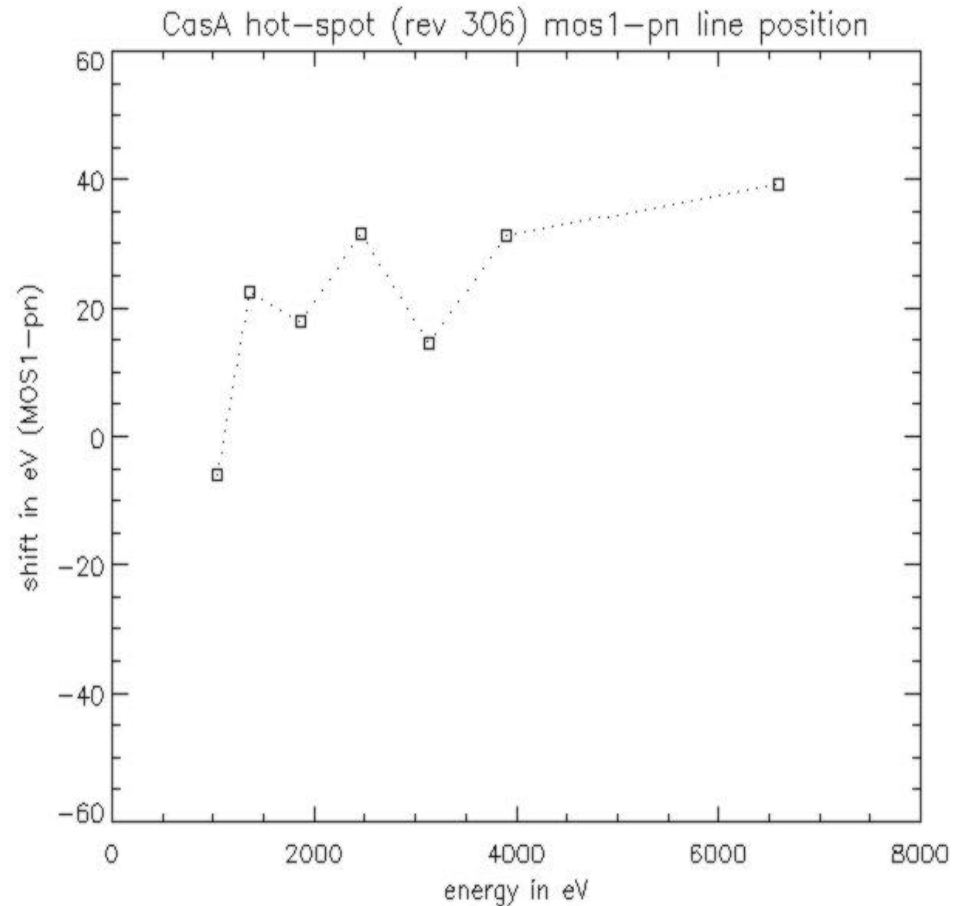


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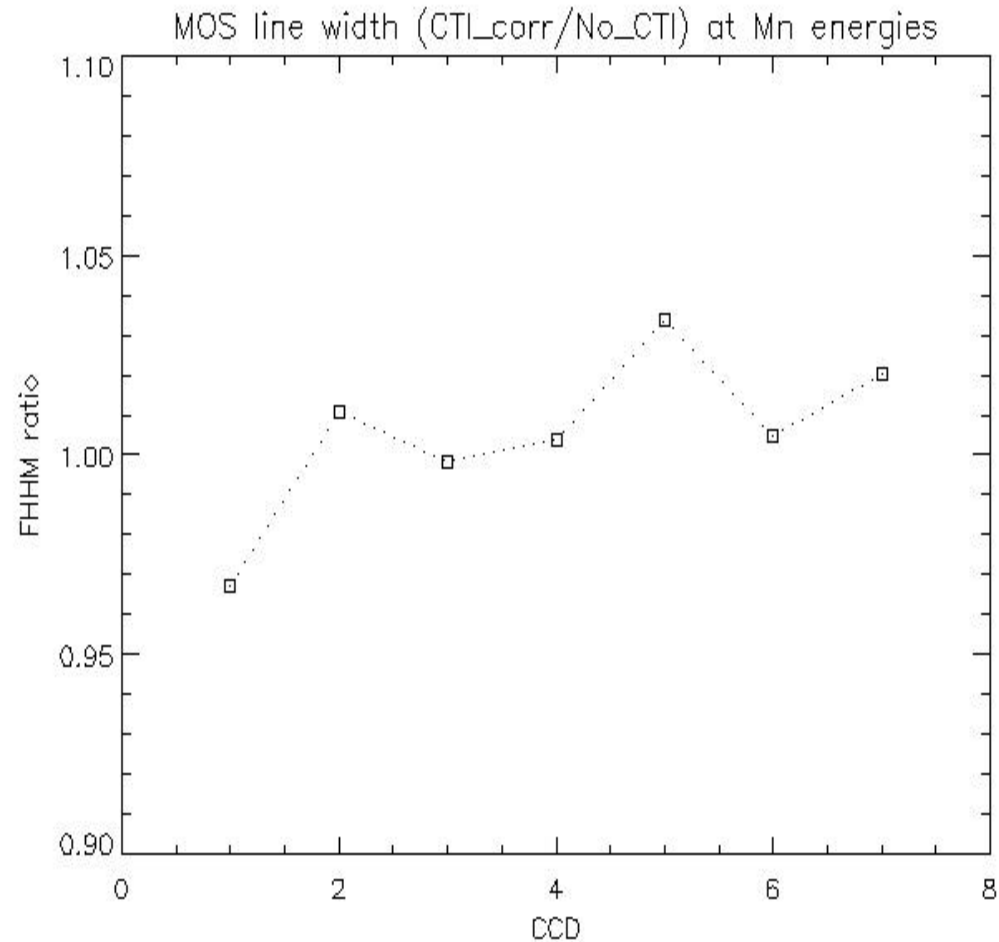
Comparison MOS/ pn

- **Comparison MOS/pn on the hot spot of CasA observed in rev 306.**
- **MOS shifted by > 20 eV compared to pn in both LW and full-frame at high energies.**
- **Note: no mode-dependent CTI correction.**
- **Future: make a systematic comparison MOS/pn/RGS on N132D for soft energies.**



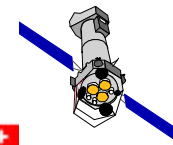
MOS energy resolution

- **Negligible line width increase after CTI correction at Mn energies**
- **$\sim < 1\%$? $\sigma = 64\text{eV}$**
- **No change or decrease at Al energies ! (35 eV)**



Conclusions

- **Status:**
 - As the linear fit CTI degradation is a good approximation, the new refined CTI algo is appropriate.
 - Give rather good results when the temperature is stabilized.
 - **Suggest to implement it in SAS v5.3**
- **Future:**
 - work on a temperature correction
 - parameters to be refined in the future with more time leverage and new CTI plots from Andrea.
 - Tune the under-correction for MOS2
 - Understand shifts with PN and compare positions with RGS at low energies (cross-calibration)



XMM-Newton

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